

# Noble Element Simulation Technique Version 2.0

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on behalf of the NEST collaboration

# NEST collaboration



UC San Diego



Stanford  
University



UNIVERSITY AT ALBANY  
State University of New York

# About NEST



- NEST (**Noble Element Simulation Technique**) is an unprecedentedly accurate and comprehensive simulation of the scintillation, ionization, and electroluminescence processes in noble elements.
- Applications:
  - Direct dark matter searches (LUX, LZ, XENON10, PandaX...)
  - Double beta decay searches (ever in a low energy ranges (nEXO, etc.)
  - Much more!

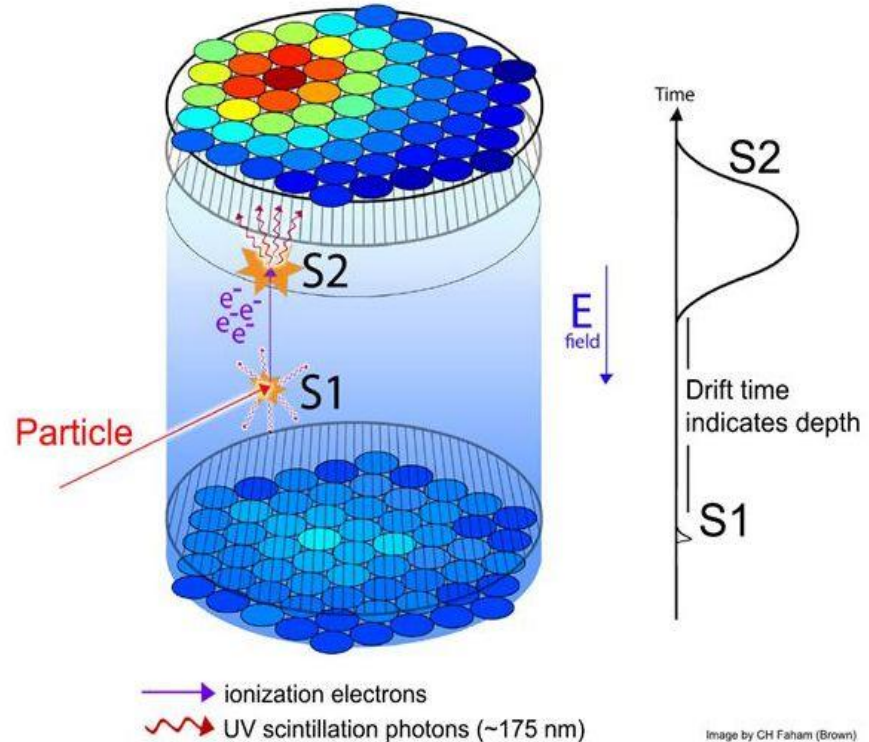


Image by CH Faham (Brown)

# What's New?

NESTv1.0	NESTv2.0
Only GEANT version	Standalone & GEANT versions
No alphas and heavy ions	Alphas and heavy ions simulations included
All equations based on theoretical models (Thomas-Imel box, Doke-Briks, etc.)	<u>Using sigmoids (family of S-shape functions)</u> , which still closely resemble those models.

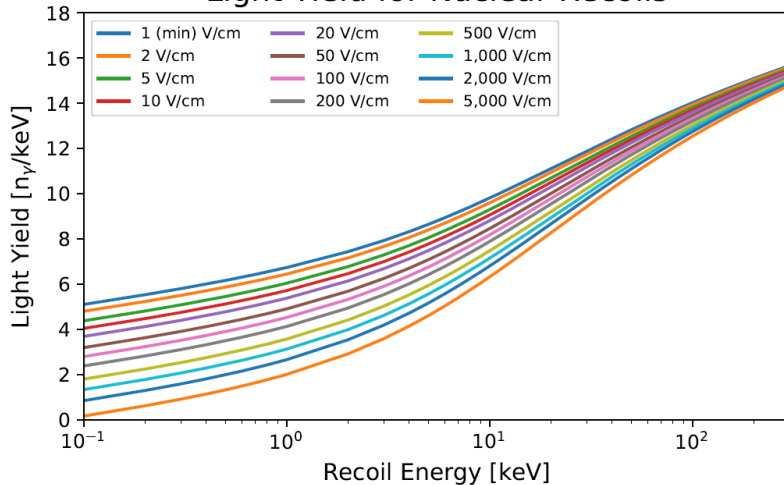
# What's New?

- Revisited old data, corrected for newer phenomena
  - Includes 2PE effect for VUV photons in PMTs
  - Allowed 'zero-field' to vary (i.e. took into account possible errors in detectors' electronics)
  - Allowed extraction efficiency to vary
  - ER:  $\beta$ -model vs.  $\gamma$ -model
- Exciton-Ion ratio is energy-dependent
- Accurately models detector effects for S1-S2 bands (means, widths, leakages)

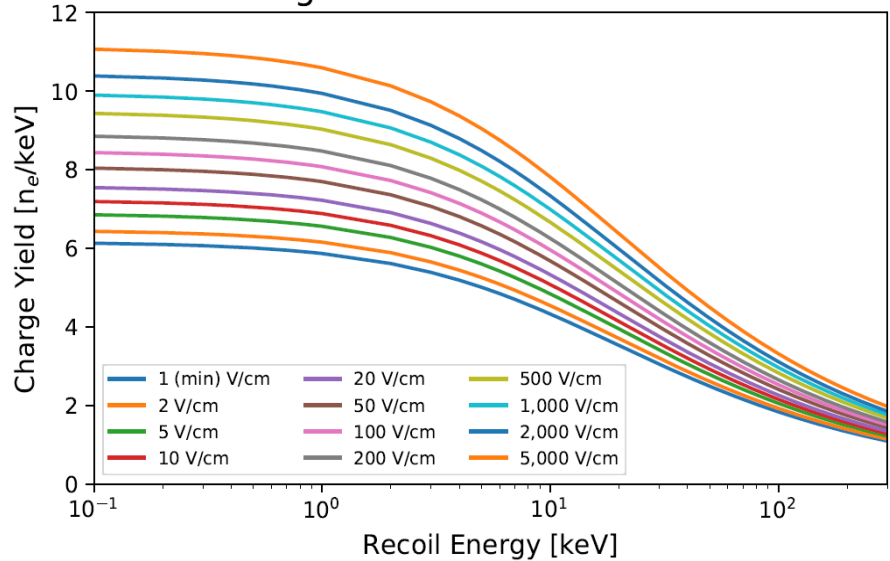
# Nuclear Recoils

- Total quanta (light+charge) is now a power law
  - $12.6 * (\text{Energy})^{1.05}$ 
    - Elegant → almost linear
  - $12.6 \pm 0.9$  &  $1.05 \pm 0.05$
- Mean-yields equations replaced with simple functions

Light Yield for Nuclear Recoils

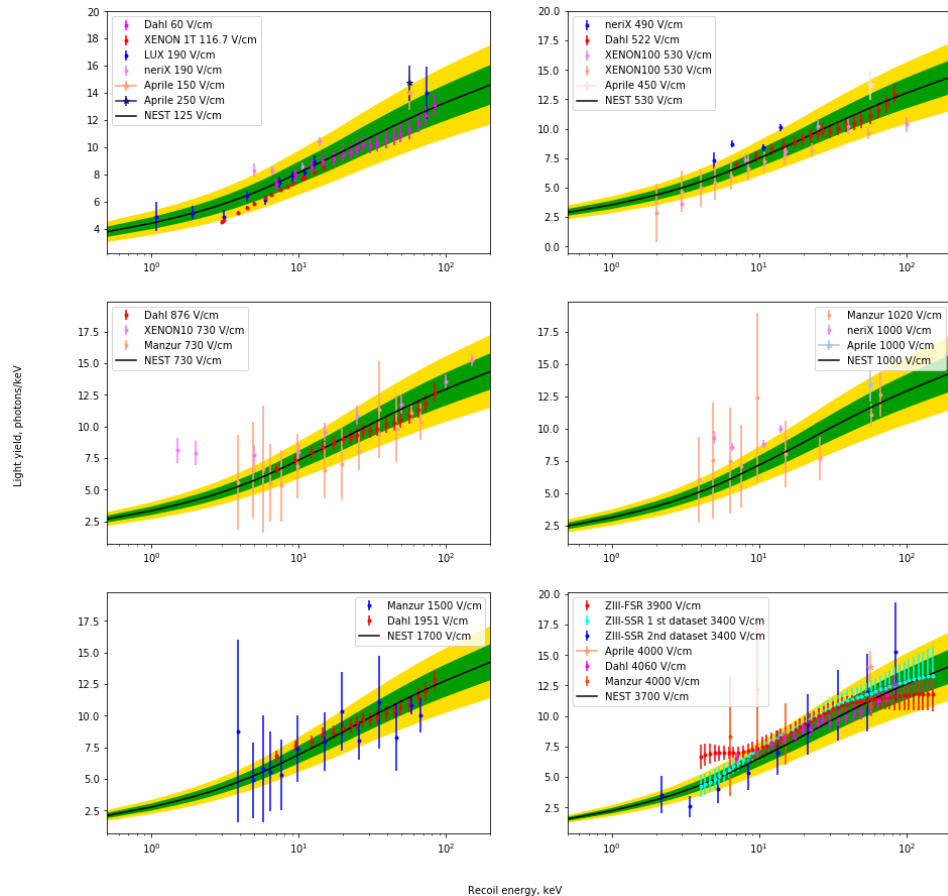


Charge Yield for Nuclear Recoils



$$Q_y = \frac{1}{A * (\text{energy} + B)^{0.5}}$$

# Nuclear Recoils

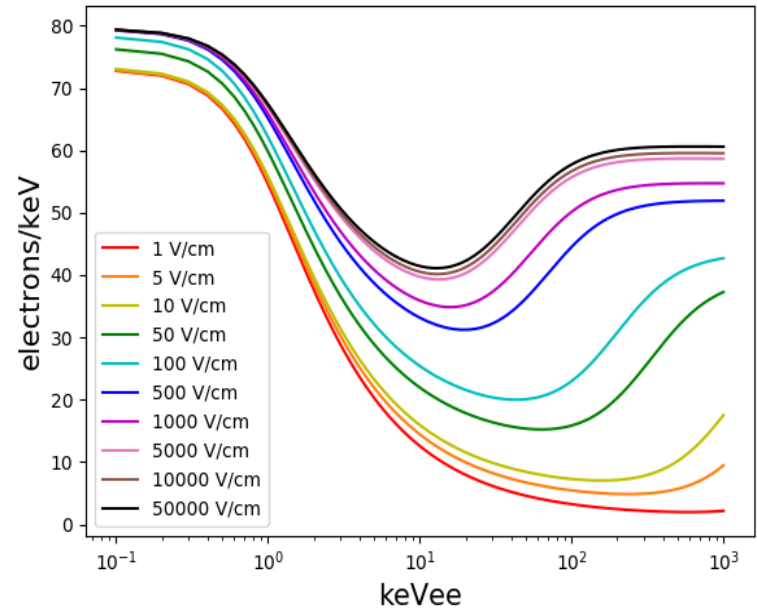
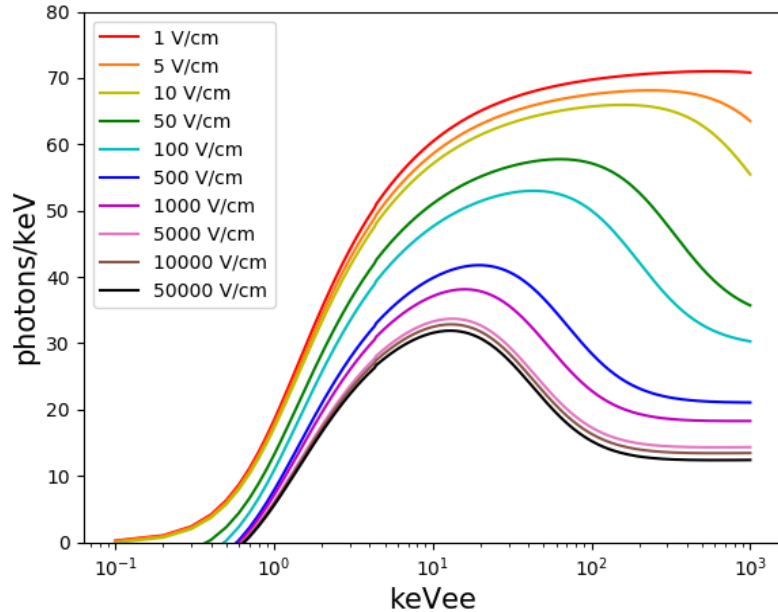


Agrees with  
known data in 1-  
sigma margin of  
error, including  
high fields  
experiments

# Electronic Recoils -- Sum of Two Sigmoids

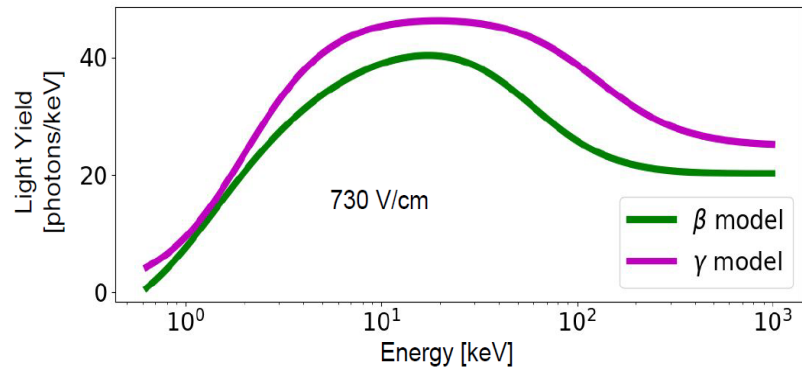
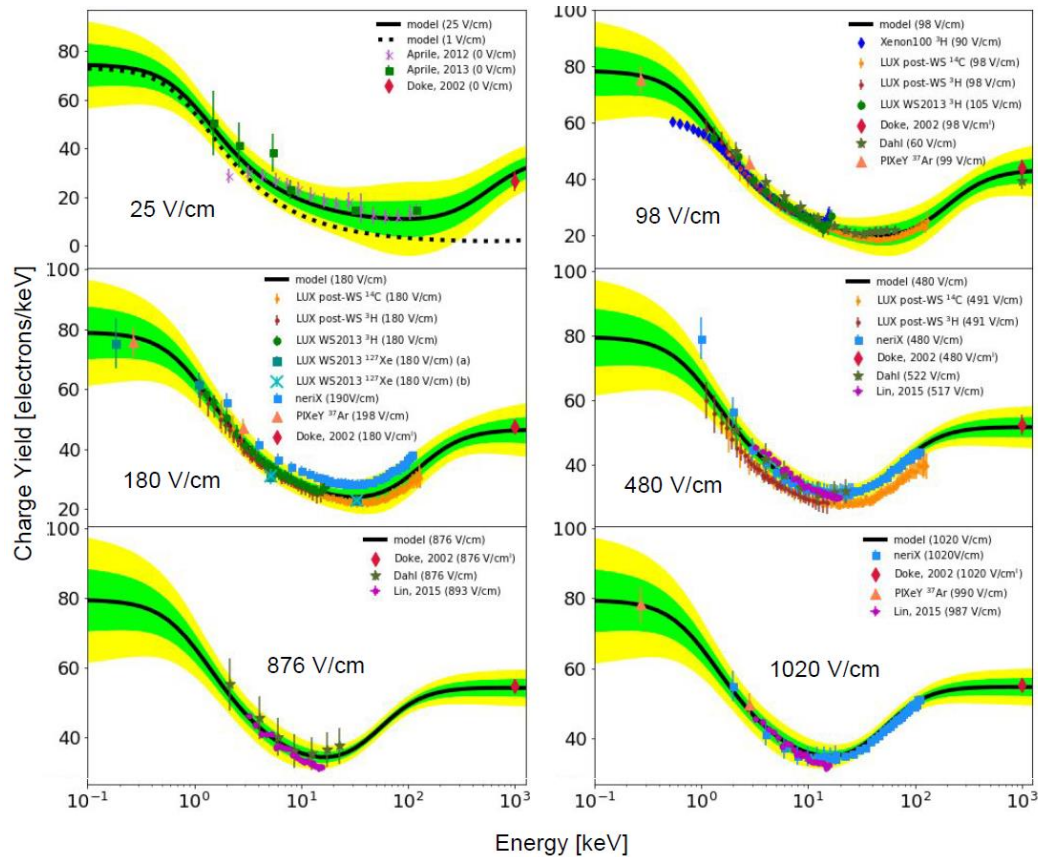
- Smooth transition between low and high energies
- $Ly + Qy = \text{const}$

Beta electron recoils: Ly and Qy





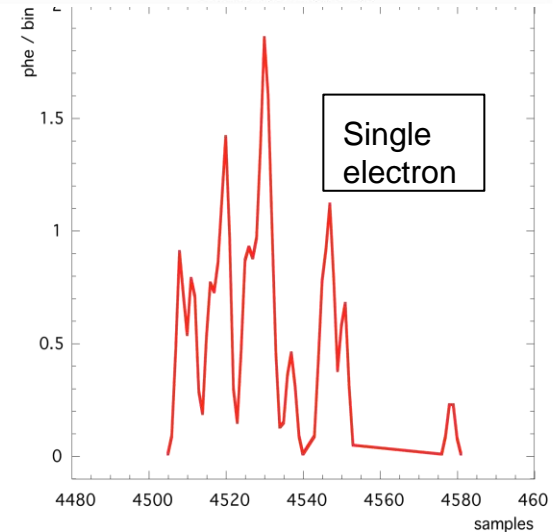
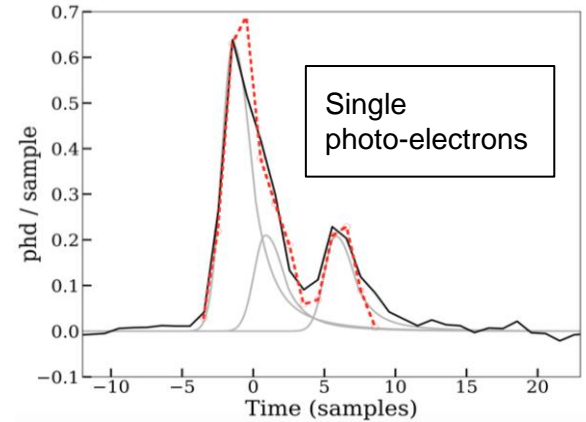
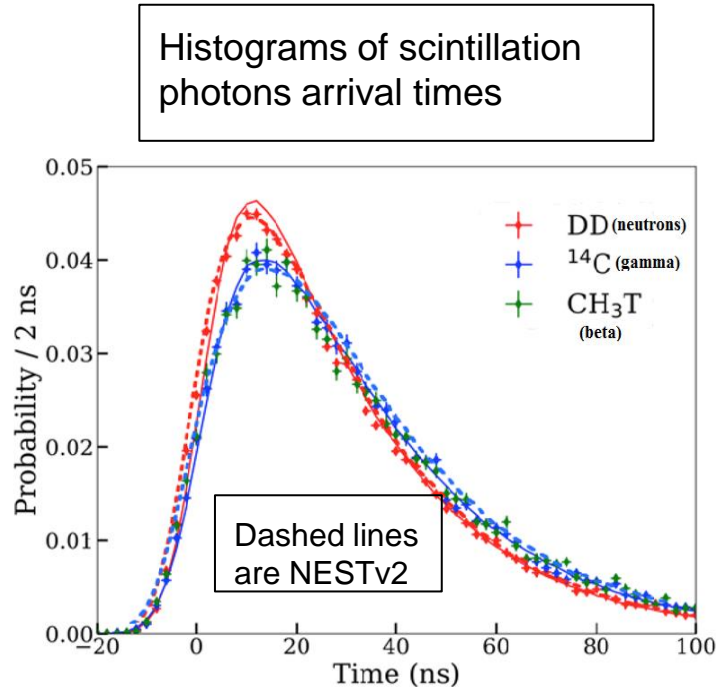
# Electronic Recoils



Two separate models for ER: beta and gamma

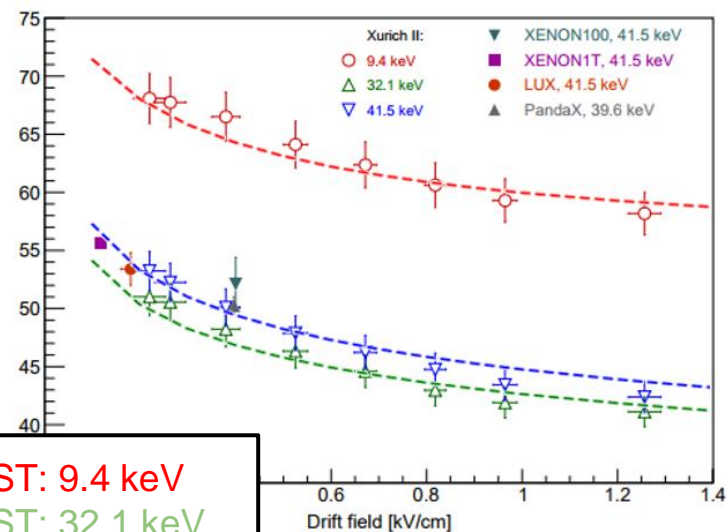
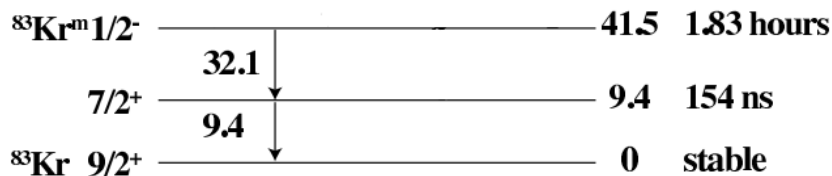
# Pulse shapes and single electrons

- Matches LUX pulse shape discrimination
- Can also simulate single electrons!
- Simulates SE noise in LXe

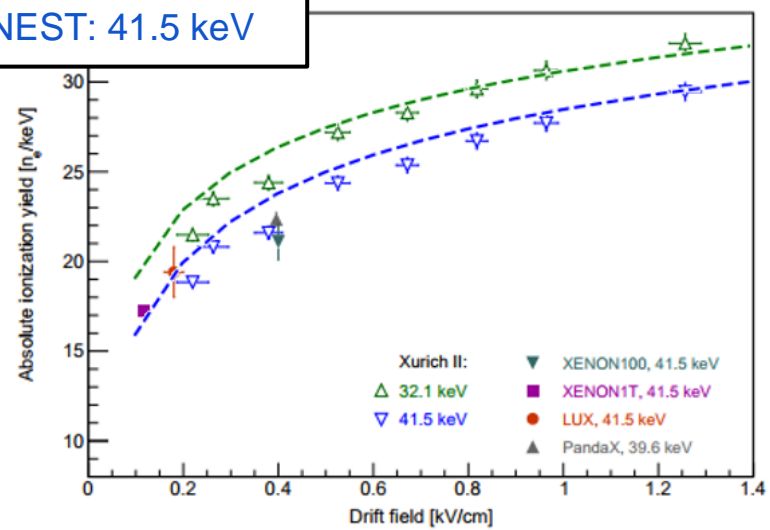


# $^{83m}\text{Kr}$

- Robust time-dependent model
- Matches individual decays as well as ‘merged’ decay



---NEST: 9.4 keV  
 ---NEST: 32.1 keV  
 ---NEST: 41.5 keV



# $^{83\text{m}}\text{Kr}$

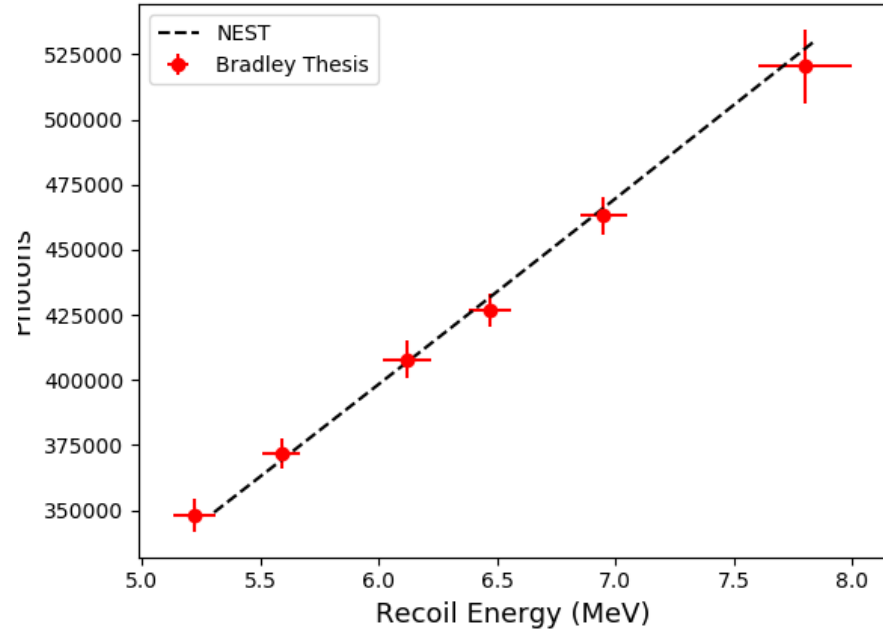
- 1  $\sigma$  agreement with LUX and XENON100

	Drift Field (V/cm)	Photons/keV, Electrons/keV	NEST Result
LUX Ly	180	$53.4 \pm 1.4$	53.0
LUX Qy	180	$19.4 \pm 1.4$	20.0
XENON100 Ly	366	$52.5 \pm 1.8$	50.6

# $\alpha$ -Model

- L-factor fixed by fitting to Adam Bradley's thesis data
  - (LUX: 180V/cm)
- Still uses Thomas-Imel box model here
  - Energy-independent for simplicity

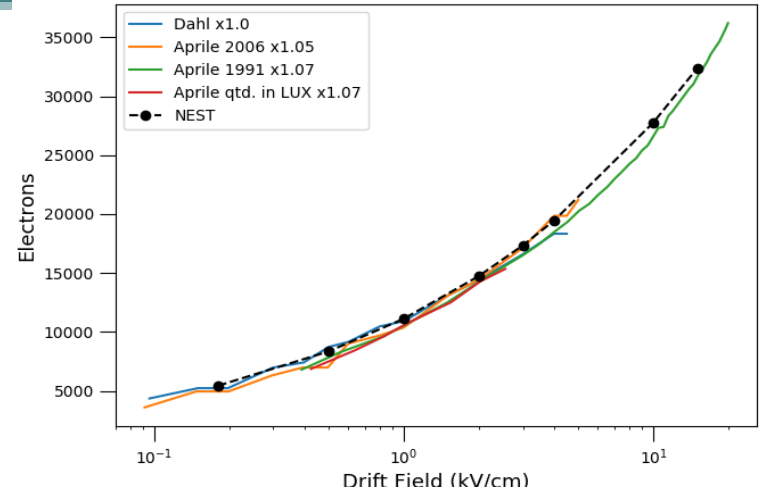
Light Yield vs. Recoil Energy from  $\alpha$ -particles



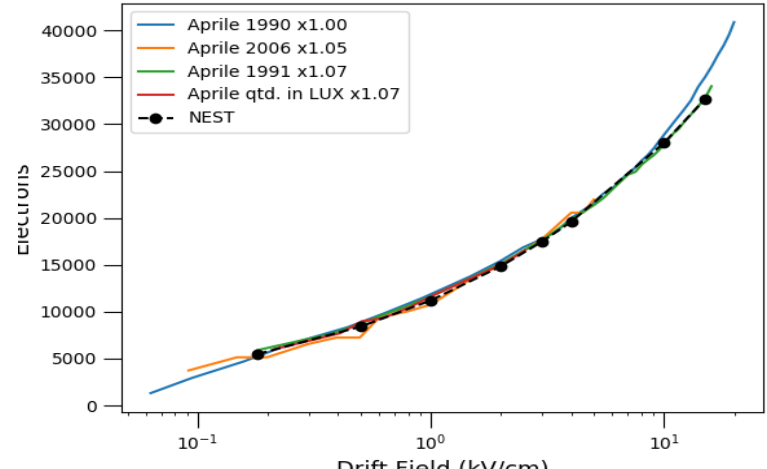
# $\alpha$ -Model

- Worked by slightly correcting data for extraction efficiency
- Good agreement for strong fields

$^{210}\text{Po}$  Charge Yields from  $\alpha$ -particles

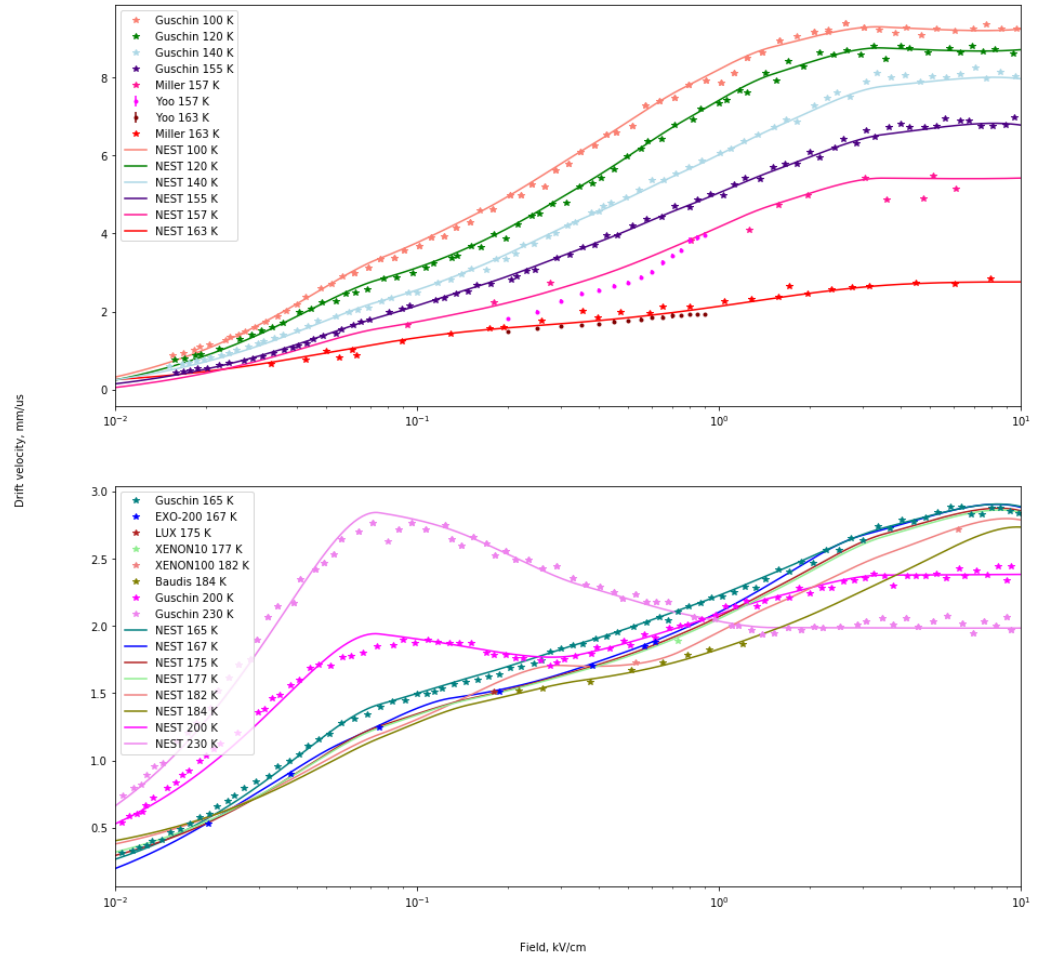


$^{241}\text{Am}$  Charge Yields from  $\alpha$ -particles



# Drift velocity

- NEST also simulates drift velocity for various xenon temperatures and states
- Has good agreement with old and new data



# Energy Resolution

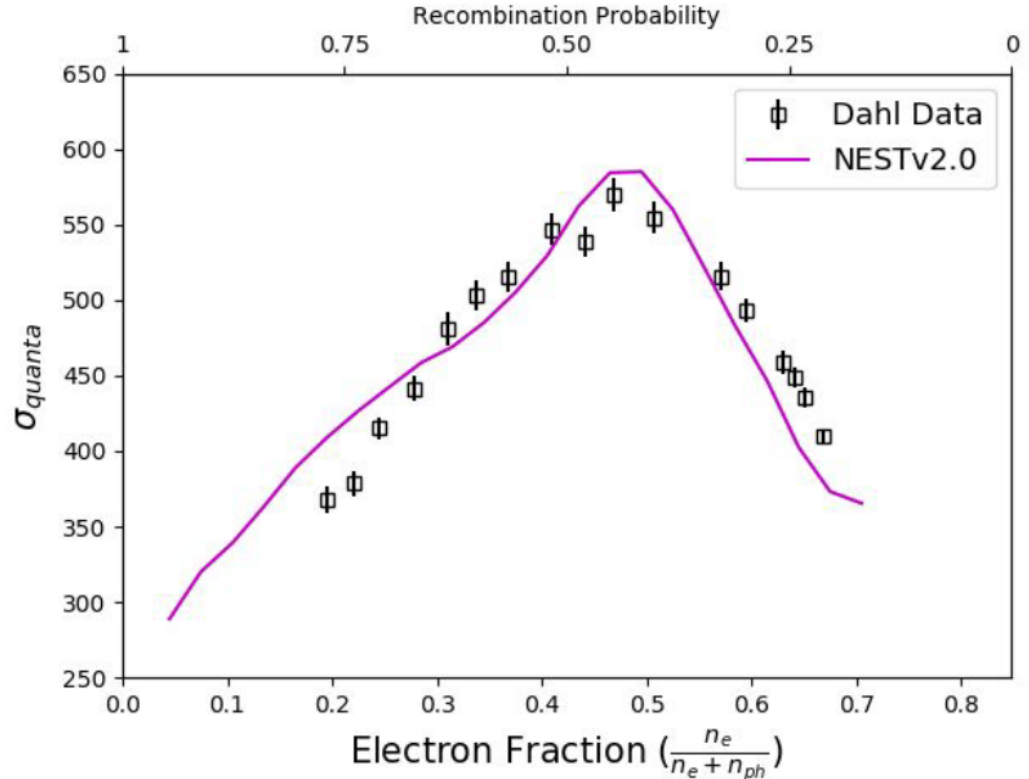
- Quantum Fluctuations
  - First estimates of fluctuations in energy resolution and fluctuations in quanta produced were by Ugo Fano in the 1940's.
  - There is energy “lost” when photons are produced in LXe from electron recoils!
  - $E = W*(n_\gamma + n_e) \rightarrow$  Work Function:  $W = 13.7$  eV
  - Fluctuations modeled using an empirical “Fano-like” factor proportional to  $\text{sqrt}(\text{energy})*\text{sqrt}(\text{field})$
- Recombination Fluctuations
  - Binomial recombination has never matched data well.
  - Same equation as cited in LUX Signal Yields Publication:  $\sigma_T^2 = (1-p)*n_i*p + (\sigma_p n_i)^2$ 
    - $\sigma_p$  in NEST is both field-dependent and energy-dependent



# Recombination Fluctuations

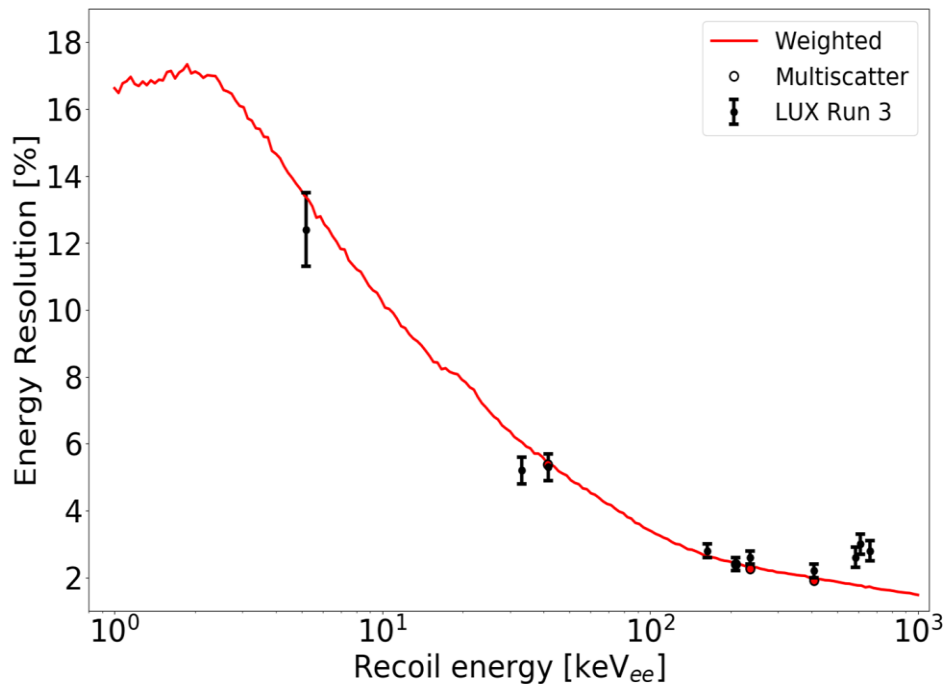
- Comparing to Eric Dahl's PhD thesis data.
- Corrected Dahl data for overestimation: corrected 15% downward for 2PE effect and extraction eff.

$$\sigma_T^2 = (1-p) \cdot n_i \cdot p + (\sigma_p n_i)^2$$

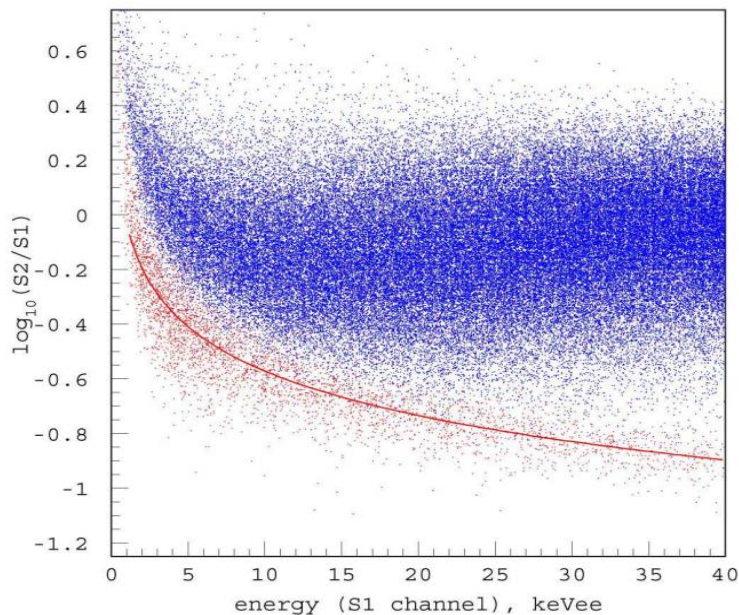


# Energy Resolution: LUX

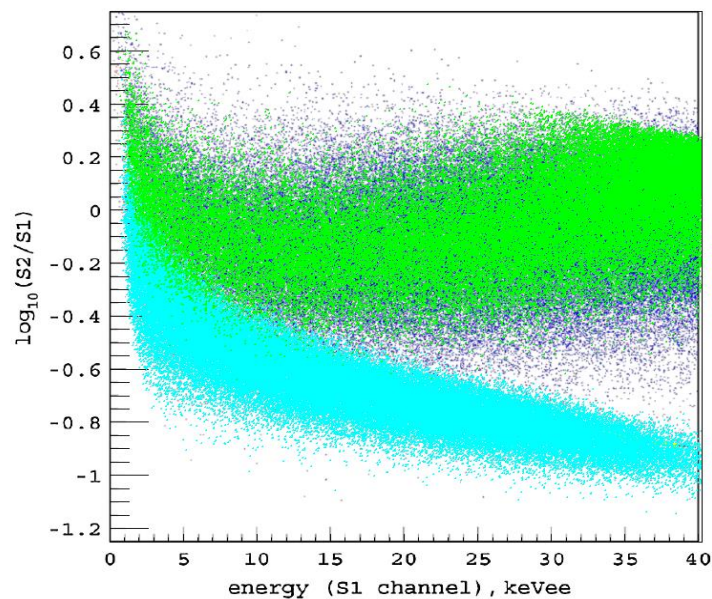
- Good fit to LUX Run 3
- $\beta$ -model better at lower energies. Fit here uses a weighted combination of NEST's  $\beta$  and  $\gamma$  models



# Energy Resolution: ZEPLIN-III



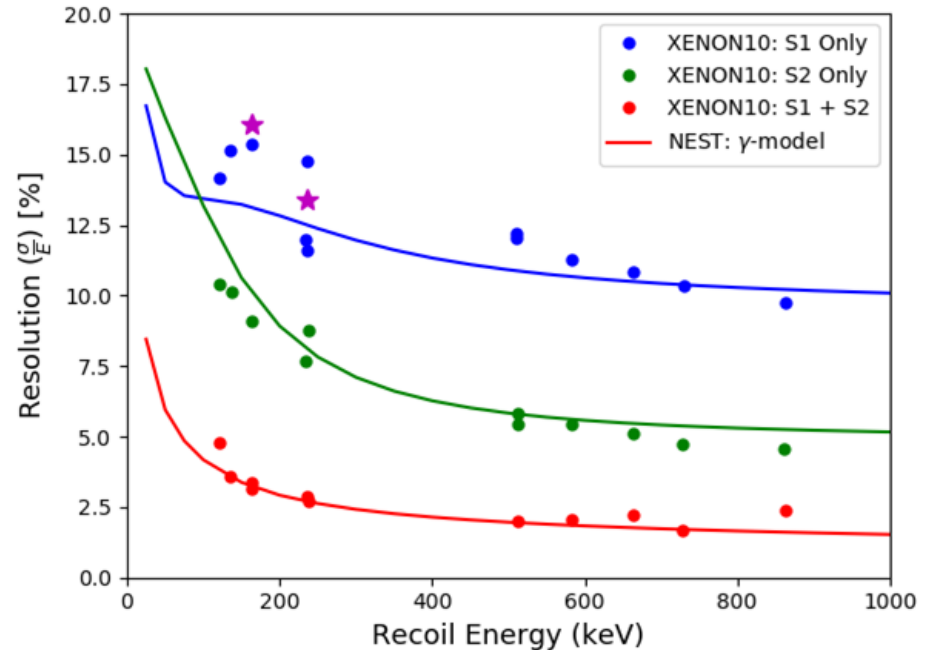
NEST matched original Z-III resolution (99.999% on AmBe and beta data) and AmBe and beta distributions (ever means)



Light blue → NEST AmBe data  
Green → NEST β data

# Energy Resolution: XENON10

- Good agreement with XENON10 energy resolution
  - Optimized a Fano-like factor for best agreement → Data suggested field & energy dependence
  - Data suggests that the Fano factor is both energy-dependent and field-dependent
- Magenta stars are  $^{129\text{m}}\text{Xe}$  &  $^{131\text{m}}\text{Xe}$ 
  - Decay in many steps, used NEST to combine the yields from each decay and added them together
  - $^{83\text{m}}\text{Kr}$  model suggests that multi-step decays have subtle time-dependence



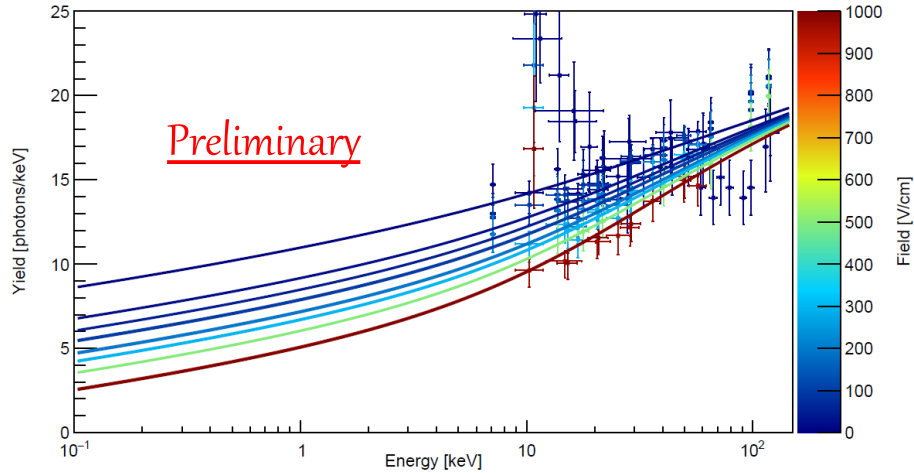
[arxiv.org/pdf/1001.2834.pdf](https://arxiv.org/pdf/1001.2834.pdf)

# Argon NEST

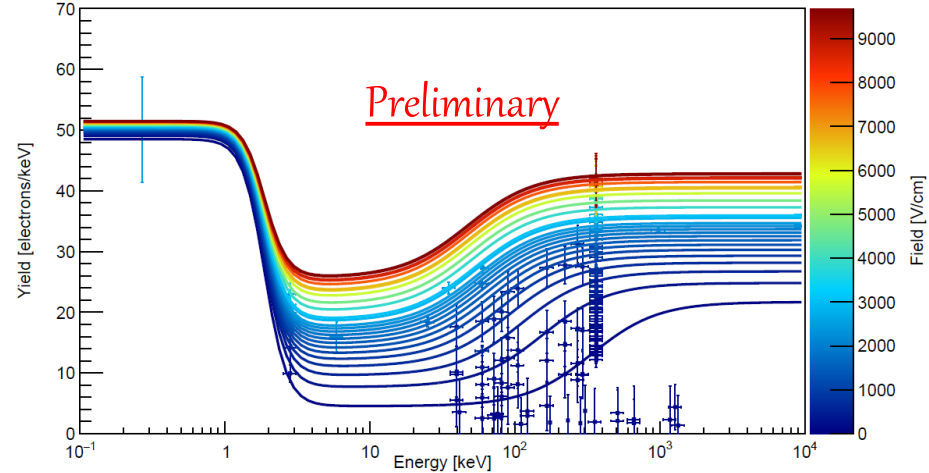
- Argon NEST is under-development version of NEST for argon
- Assumption: both Xe and Ar are noble elements – and formulae would be similar sigmoids for Ar too
- Empirical models for argon are very important – because theoretical models sometimes are contradictory to each other

# Argon NEST

NR Light Yield at Various Fields



ER Charge Yield at Various Fields



- Preliminary models for NR and ER are ready
- Kr and heavy ions models are under development

# Conclusion

- NESTv2 is a powerful simulation tool, which now has two versions: standalone tool and GEANT4 library.
- Accurately simulates many different interactions in LXe and GXe (argon models currently in process)
- User-friendly code so you can add any other interactions that you might find useful.
- Get yourself a copy!
  - <https://github.com/NESTCollaboration/nest>
  - [nest.physics.ucdavis.edu](http://nest.physics.ucdavis.edu)

**Thank you for your attention!**



# Backup Slides

# Heavy Nuclei

- Expanded the  $\alpha$ -model to include scattering events with heavy ions
- Again, contradictory data sets, splits the difference

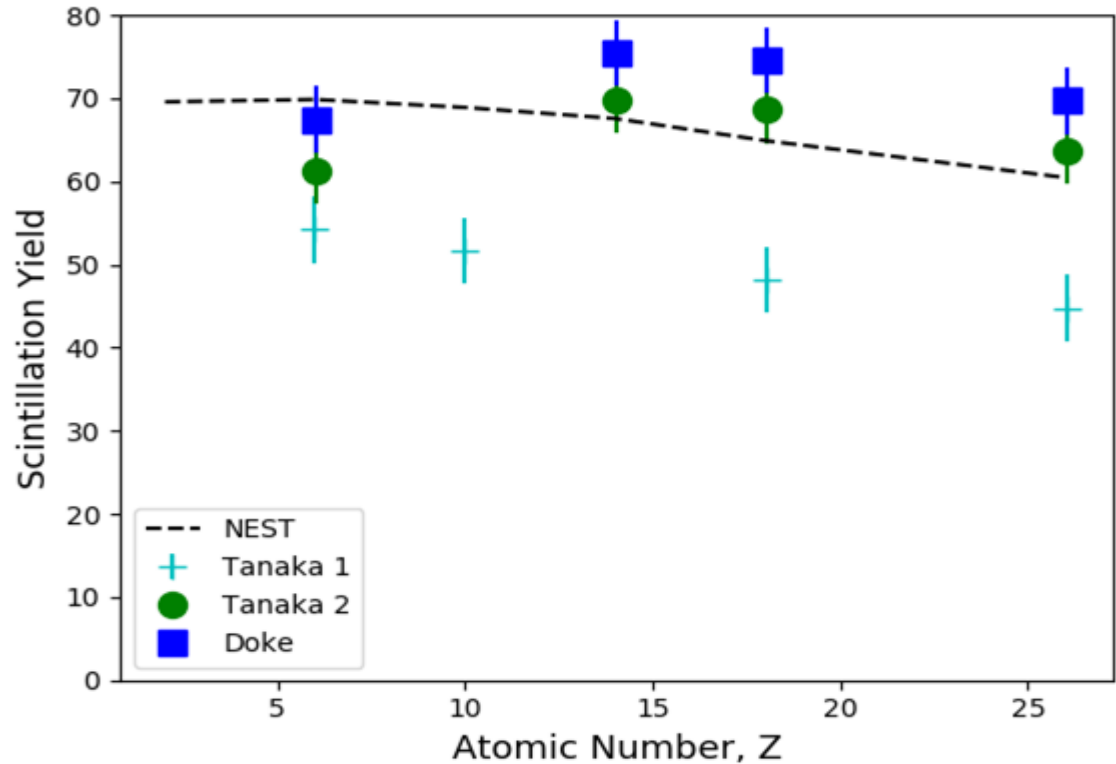
## Absolute Scintillation Yields in Liquid Argon and Xenon for Various Particles

T. Doke, et. al. 2002.

[Japanese Journal of Applied Physics, Volume 41, Part 1, Number 3A](#)

## LET dependence of scintillation yields in liquid xenon

M. Tanaka, et. al. 2001



# Boron-8

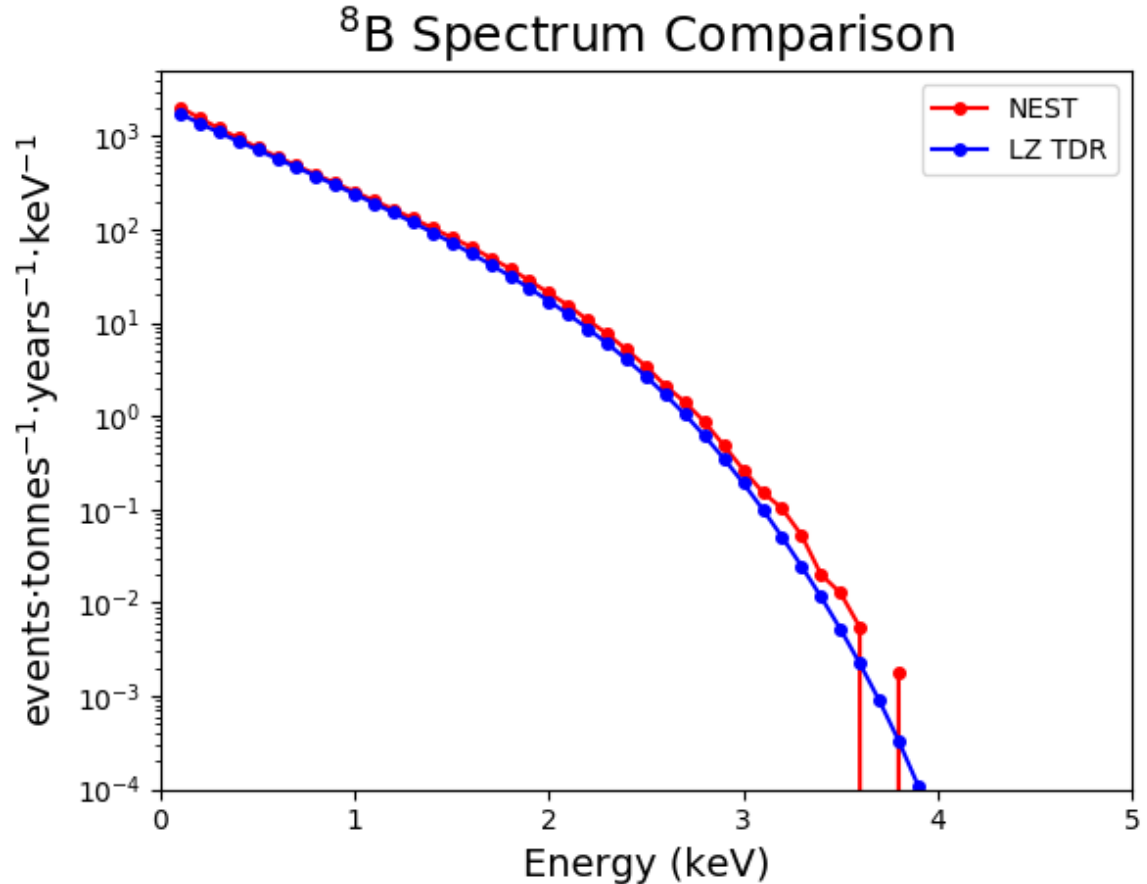
- Great agreement with LZ TDR  $^8\text{B}$  spectrum.
- Not a look-up table!

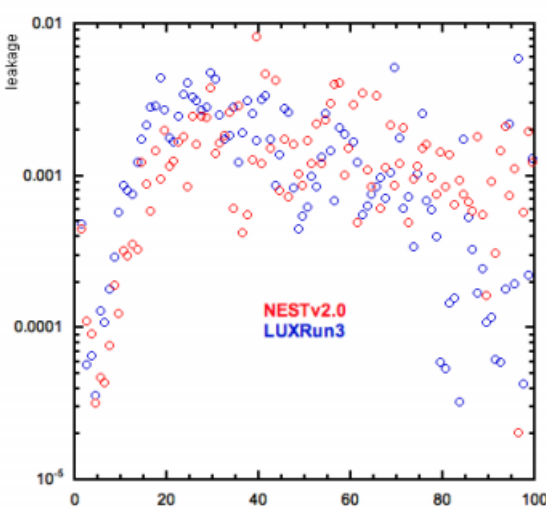
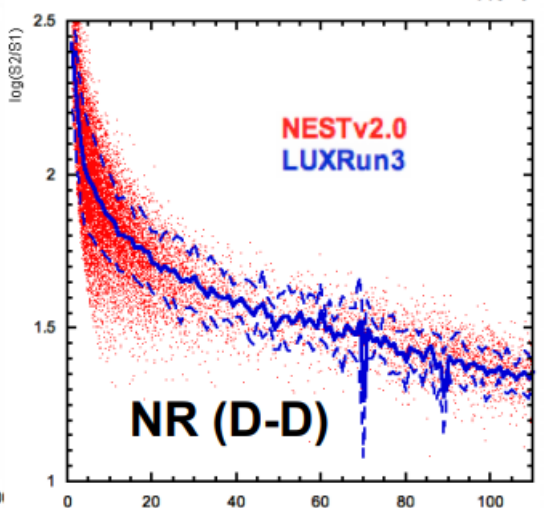
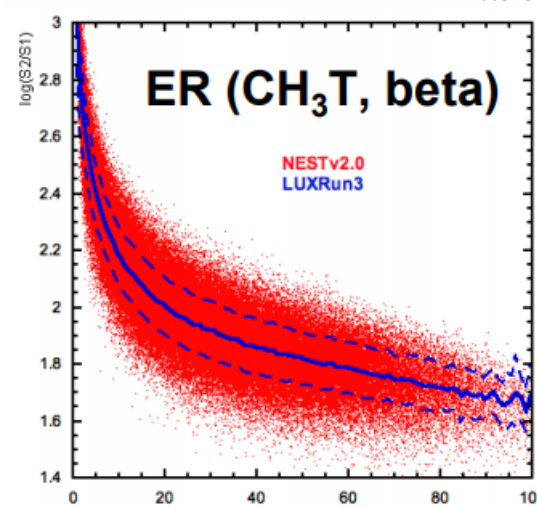
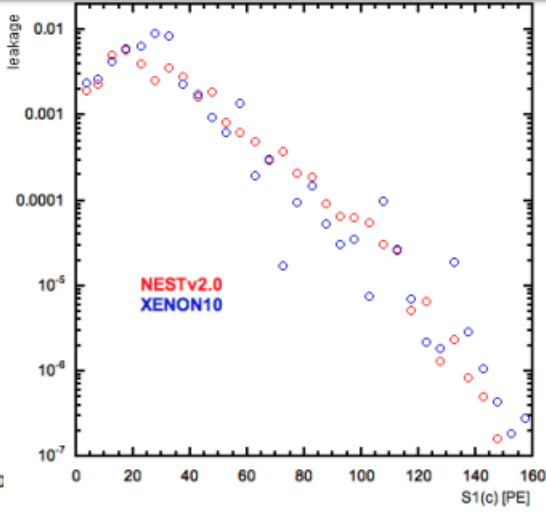
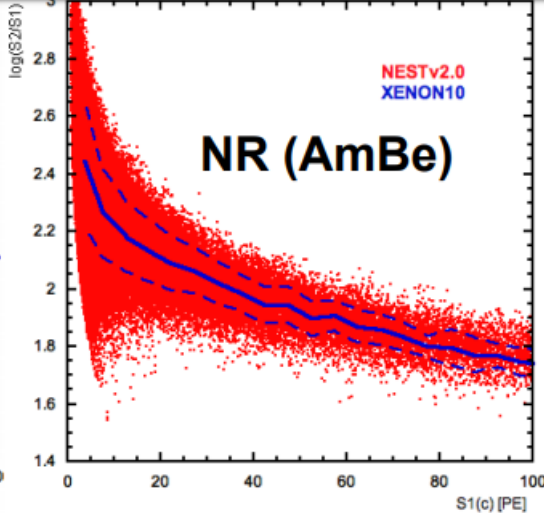
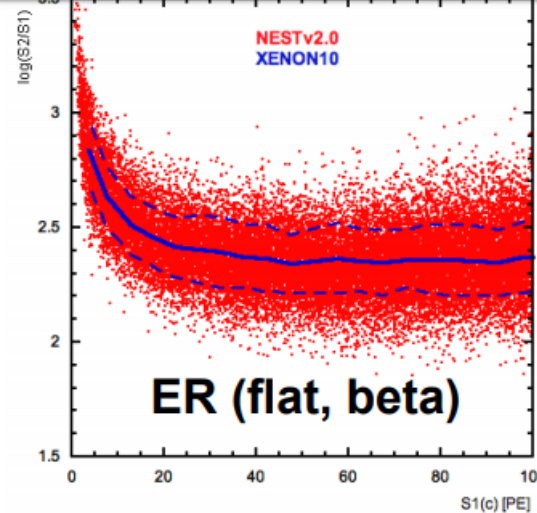
[LUX-ZEPLIN \(LZ\) Technical Design Report](#)

[B.J. Mount \(Black Hills State U.\) et al.](#), Mar 27, 2017.  
392 pp.

LBNL-1007256, FERMILAB-TM-2653-AE-E-PPD

e-Print: [arXiv:1703.09144](#) [physics.ins-det]





Going Full-Scale with Detector Effects

NESTv2 accurately reproduces NR and ER bands.

X10 detector file example provided with NESTv2. Easy to change and add your own detector!